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Enhancement of stratospheric aerosols after solar proton event

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Abstract. The lidar measurements at Verhnetulomski observatory (68.6°N, 31.8°E) at Kola peninsula detected a considerable increase of stratospheric aerosol concentration after the solar proton event of GLE (ground level event) type on the 16/02/84. This increase was located at precisely the same altitude range where the energetic solar protons lost their energy in the atmosphere. The aerosol layer formed precipitated quickly (1–2 km per day) during 18, 19, and 20 February 1984, and the increase of $R(H)$ (backscattering ratio) at 17 km altitude reached 40% on 20/02/84. We present the model calculation of CN (condensation nuclei) altitude distribution on the basis of an ion-nucleation mechanism, taking into account the experimental energy distribution of incident solar protons. The meteorological situation during the event was also investigated.

1 Introduction

It is now well documented that heterogeneous chemistry triggering ozone destruction occurs on sulfate aerosols (Tolbert *et al.*, 1988; Hofman and Solomon, 1989). The main sources of sulfate aerosols in atmosphere are volcanic eruptions, anthropogenic emissions and biological activity. Contribution of each component to aerosol layer formation in the stratosphere differs in the two hemispheres and depends on latitude. The anthropogenic pollution seems to be the major contributor of sulfuric acid in Arctic region compared with Antarctic where the phytoplankton emissions seem to prevail (Langner *et al.*, 1992). However, volcanos also sometimes play a dominant role as sources of sulfur injections into the stratosphere. For example, the eruption of Mount Pinatubo (1991) caused increases in sulfuric acid abundance by a factor of 50 (Tolbert, 1994). Another source of stratospheric aerosols

exists in nature besides those already mentioned. The mechanism of ion nucleation triggering by solar cosmic ray ionization was suggested by Arnold (1982) as a potential source of condensation nuclei (CN) on which stratospheric sulfate aerosols can form (Arnold, 1982; Hofman and Rosen, 1983).

Ion nucleation involves growth of ions by association of sulfuric acid molecules to a critical size at which the resulting molecular cluster is stable against ion-ion recombination. The rate of formation of nuclei is proportional to the ion production rate, and thus temporal and spatial changes of atmospheric ionization due to solar and galactic cosmic rays may induce corresponding changes of condensation nuclei concentration.

The creation of sulfate aerosols by means of ion nucleation mechanism seems to be a possible reason for the ozone “minihole” appearance resulting from a set of heterogeneous chemical reactions after solar proton events of GLE type in May 1990 at Spitsbergen (Shumilov *et al.*, 1992; 1995). In this work the mechanism of aerosol formation during the GLE event on 16 February 1984 is discussed.

2 Experimental results

The solar proton event (SPE) of 16 February 1984 was detected at 08:58 UT and was caused by a solar flare behind the W-limb of the Sun. It had a very rigid energy spectrum and belonged to the GLE type of SPE. The maximum intensity of solar protons during the GLE reached $5 \text{ particles cm}^{-2} \text{ s}^{-1} \text{ ster}^{-1}$ for energies between 84 and 200 MeV (Hargreaves *et al.*, 1987). Figure 1 shows the Apatity ($\Phi' = 63.3^\circ$) neutron monitor data for February 1984.

During February 1984 the stationary lidar operated at Verhnetulomski observatory (68.6°N, 31.8°E) located in the Murmansk region. At the lidar the diameters of receiving and transmitting telescopes were 50 cm and 20 cm, respectively. Two lasers with emission wavelengths of 694.3 nm and 532.0 nm were used in the lidar. All profiles

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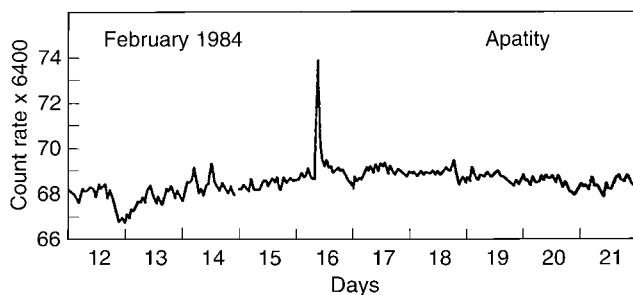


Fig. 1. Hourly count rates of the neutron monitor at Apatity observatory from 12 to 21 February 1984

shown here relate to the 694.3 nm channel. The ruby laser has wavelength of 694.3 nm, pulse repetition – 1 Hz, energy per pulse – 0.4 J and pulse width – 30 ns. To limit the background emission an interference filter with 1.8 nm bandwidth and 0.4 transmission coefficient is used in the receiver. Lidar permits measurements in an altitude range between 3 and 40 km with vertical resolution of 1.5 km.

In Fig. 2 profiles of aerosol backscatter ratio $R(H)$ at wavelength of 694.3 nm are shown. The expression for aerosol backscatter ratio is as follows:

$$R(H) = (\beta_{mol} + \beta_{aer})/\beta_{mol}, \quad (1)$$

where β_{mol} and β_{aer} are molecular and aerosol backscattering respectively. Figure 2 shows 15, 18, 19, and 20 February $R(H)$ profiles relative to the curve of average values for previous five undisturbed days. The lidar measurements on 16 and 17 February were absent because of bad weather conditions. The aerosol backscatter increase on 18 February 1984 between 15 and 25 km altitude can be seen (Fig. 2).

To examine the influence of meteorological factors an analysis of the meteorological situation over the Kola peninsula region has been made. The 100-hPa meteorological map analysis has shown that during the period from 13 February to 20 February, the Kola peninsula was in the coldest part of the circumpolar vortex. Some altitude temperature profiles are shown in Fig. 3. These profiles were obtained at Murmansk meteorological observatory (68.3°N, 33°E) located close to Verhnetulomski observatory. Figure 3 clearly shows that the temperatures are low enough for PSC formation.

Additional information about the meteorological situation before and after GLE event considered may be obtained from Fig. 4 where the 100-hPa (16-km altitude) temperatures for Kola peninsula during the second half of February 1984 are shown. The arrows indicate the temperature at which Type-1 PSCs can be formed (Nagatani *et al.*, 1990). Figure 4 demonstrates that the 100-hPa temperatures were sufficiently low, and PSC activity could be expected over Kola peninsula between 13 February and 21 February. However, the aerosol enhancement seems not to be related to the temperature decrease on 16 February (–84 °C). The same extreme cold temperature was observed on 13 February (–83 °C) without any aerosol backscatter ratio $R(H)$ changes (for reference see Figs. 2, 3). It seems reasonable to say that the PSC could form even before 13 February, and the additional temper-

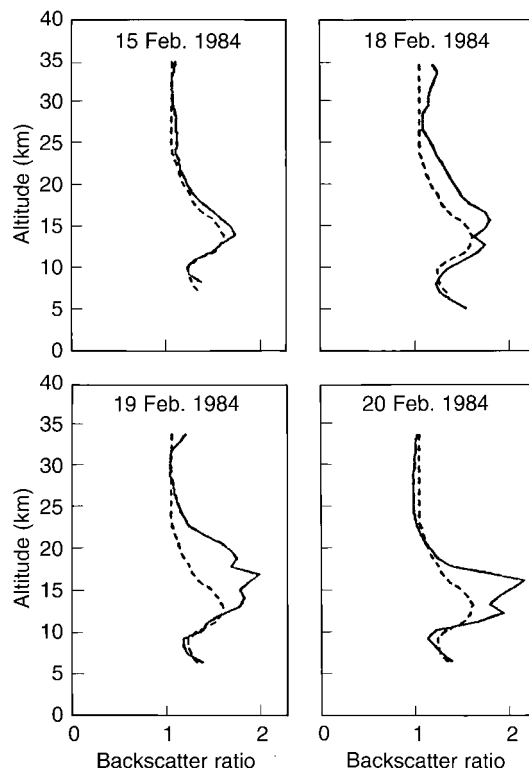


Fig. 2. The vertical profiles of scattering ratio $R(H)$, $\lambda = 694.3$ nm taken on 15, 18, 19, and 20 February 1984 at Verhnetulomski observatory. Dashed curves give the quiet day level averaged for previous five undisturbed days

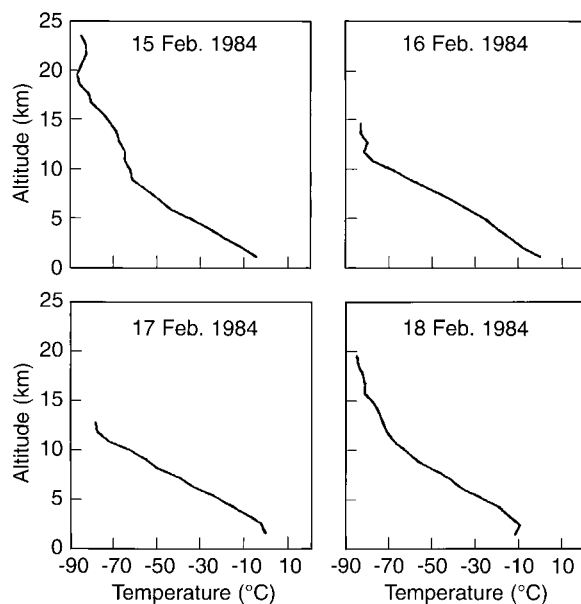


Fig. 3. Temperature profiles measured by balloons on 15, 16, 17, and 18 February 1984 above Murmansk observatory

ature decrease observed on 16 February could not lead to PSC formation. We therefore assume the aerosol increase observed between 15 km and 20 km was probably caused by a GLE event.

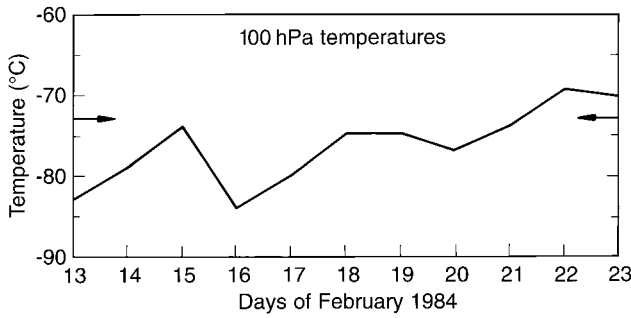


Fig. 4. 100-hPa temperatures over Kola peninsula from 13 February to 23 February 1984. Arrows indicate the threshold temperatures for type 1 PSC formation

3 Data analysis and discussion

Figure 2 shows aerosol backscatter increase on 18 February 1984 relative to the average profile for the previous 5 days. The increase was not associated with considerable temperature enhancements (see Fig. 3) and seems to be caused by the GLE event on 16 February 1984.

Taking into consideration that $\beta_{mol} = n_m \sigma_m$ and $\beta_{aer} = n_a \sigma_a$, where σ_m and σ_a are molecular and aerosol backscattering cross sections, and n_m , n_a are molecular and aerosol concentrations, respectively, Eq. (1) can be rewritten as:

$$R - 1 = (\sigma_a n_a) / (\sigma_m n_m). \quad (2)$$

Aerosol concentration can be derived from backscattering ratio at 694.3 nm ($R_{0.69}$). A reasonable aerosol backscattering cross section $\sigma_a = 2.5 \cdot 10^{-11} \text{ cm}^2 \text{ ster}^{-1}$ and molecular backscattering cross section $\sigma_m = 2 \cdot 10^{-28} \text{ cm}^2 \text{ ster}^{-1}$ at the 694.3 nm wavelength are taken from DeLuisi *et al.* (1975) and Kent and Wright (1970) respectively. Substituting σ_a and σ_m and assuming them to be independent of altitude, from Eq. (2) we obtain:

$$n_a = 0.8 \cdot 10^{-17} (R_{0.69} - 1) n_m. \quad (3)$$

Using the backscatter ratio profile $R_{0.69}(H)$ and a simple model of the atmosphere (Russel *et al.*, 1979), we have calculated from Eq. (3) the height profile of aerosol concentration.

In Fig. 5 the experimental aerosol background profile (*Nb*) and disturbed aerosol profile (*Nd*) of 18 February 1984 are shown. These experimental values correspond to ordinary altitude distribution of aerosol particles with sizes $r \geq 0.69 \mu\text{m}$ (Hofman *et al.*, 1983). Calculated aerosol densities have errors limited by 4% (DeLuisi *et al.*, 1975; Bajdalo, 1978). Figure 5 shows that aerosol concentration increase on 18 February 1984 was about 50%.

To calculate ion production rate $Q(\text{cm}^{-3} \text{ s}^{-1})$ we have used the following equation:

$$Q(H) = (2\pi/\varepsilon) \cdot \int_{E_1}^{\infty} D(E) E_{Li}(E) dE, \quad (4)$$

where $E_{Li}(E)$ is the energy loss in the i th altitude slab in MeV (Jackman *et al.*, 1980), and $\varepsilon = 35 \text{ eV}$ is the energy

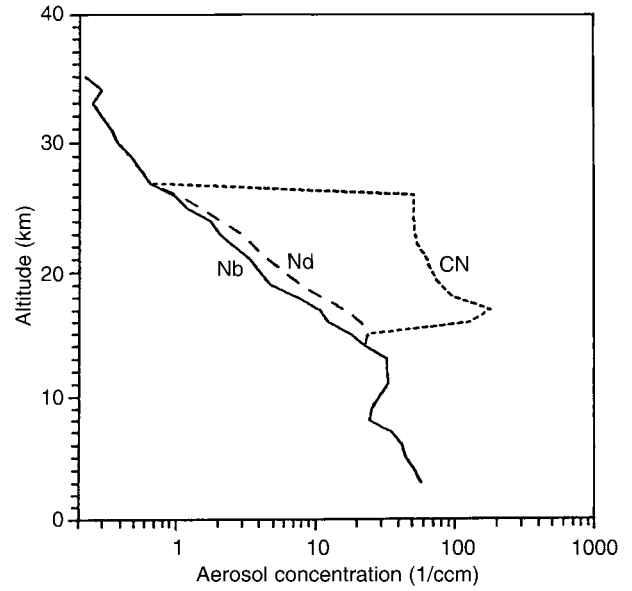


Fig. 5. Aerosol concentration profiles calculated from experimental lidar measurements on 15 February 1984 (*Nb*) is quiet day level, and on 18 February 1984 (*Nd*) is disturbed level. Curve *CN* (condensation nuclei) gives *CN* concentration profile calculated for GLE event using ion nucleation mechanism

needed for one ion-electron pair production. $D(E)$ is the differential energy spectrum of incident protons ($\text{cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1} \text{ MeV}^{-1}$) taken from Hargreaves *et al.* (1987), and E_1 is the energy of Murmansk observatory geomagnetic cut-off calculated as 150 MeV approximately.

Figure 6 shows the ion production rate profile calculated at the maximum intensity on 16 February 1984 GLE event for $E_1 = 150 \text{ MeV}$ (solid curve) and for $E_1 = 0$ (dashed curve). $E_1 = 0$ for geomagnetic polar cap region because the solar protons at all energy levels have free access to an atmosphere there. From Fig. 6 it is clear that these two curves have a considerable difference only above 30 km altitude.

Ion nucleation rate $J(\text{cm}^{-3} \text{ s}^{-1})$ is related to ion production rate Q by the equation (Hofman and Rosen, 1983):

$$J = Q \cdot (1 + (\alpha Q)^{1/2} / K [\text{H}_2\text{SO}_4])^{-N_c}, \quad (5)$$

where α is the ion recombination coefficient ($10^{-7} \text{ cm}^3 \text{ s}^{-1}$), K is the association rate coefficient ($10^{-9} \text{ cm}^2 \text{ s}^{-1}$), and $[\text{H}_2\text{SO}_4]$ is the molecular concentration of H_2SO_4 . N_c is the number of H_2SO_4 molecules in a critically sized embryo. In stratospheric conditions only molecules of H_2SO_4 and water can form stable clusters, and N_c is 3 ($h \leq 27 \text{ km}$) and 10 ($h > 27 \text{ km}$) respectively (Arnold, 1982).

Figure 5 shows the condensation nuclei (*CN*) profile calculated for the 16 February 1984 GLE event using Eq. (5) and altitude distribution of H_2SO_4 for winter conditions (Arnold, 1982). As we can see from the Fig. 5, calculated values of *CN* are one order greater than experimental ones of *Nd*. Note that high and low levels of *CN* increase coincide with experimental ones. The high level ($h = 27 \text{ km}$) corresponds to the altitude of H_2SO_4 maximum supersaturation (Arnold, 1982); the low one

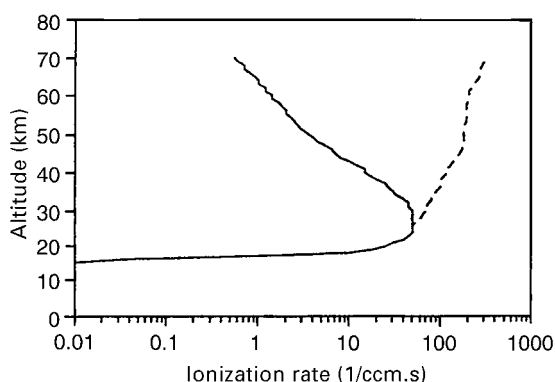


Fig. 6. Ionization rate profiles calculated for GLE event of 16 February 1984 at time of its maximum. *Solid curve* gives ionization rate profile calculated with geomagnetic cut-off E_1 at 150 MeV, and *dashed curve* gives ionization rate profile calculated when $E_1 = 0$

($h = 16$ km) is related to ionization rate decrease (see Fig. 6). But we have to take into account the fact that aerosol particles detected by lidar have radii of more than $0.69 \mu\text{m}$ compared with smaller radii of CN at $0.01 \mu\text{m}$ (Arnold, 1982). These radii differences can be reduced by involving coagulation process in terms of a kinetic equation (Twomey, 1977):

$$dN/dt = 0.5 \int_0^\infty \int_0^\infty K(X_1, X_2) n(X_1) n(X_2) dX_1 dX_2, \quad (6)$$

where N is the total number of particles, $n(X_1)$ and $n(X_2)$ are the numbers of particles of sizes X_1 and X_2 respectively, and $K(X_1, X_2)$ is the coagulation coefficient. In our case we have CN where radius is $0.01 \mu\text{m}$ and aerosol particles where radius is more than $0.69 \mu\text{m}$. So, Eq. (6) may be rewritten as:

$$dN/dt = -0.5K_1 n^2 - K_2 nNb, \quad (7)$$

where n is the number of calculated CN of radius $0.01 \mu\text{m}$, and Nb is the number of measured background aerosol particles of radius more than $0.69 \mu\text{m}$. The coefficient K_1 ($24 \cdot 10^{-10} \text{ cm}^3 \text{ s}^{-1}$) accounts for self-coagulation of CN, and the coefficient K_2 ($78000 \cdot 10^{-10} \text{ cm}^3 \text{ s}^{-1}$) accounts for coagulation of CN with large aerosol particles with radii exceeding $0.69 \mu\text{m}$ (Twomey, 1977).

Using Eq. (7) and assuming K_1 and K_2 to be constant we have calculated the time necessary for CN to transform from tiny particles whose radius is $0.01 \mu\text{m}$ to large aerosols with radii exceeding $0.69 \mu\text{m}$ through coagulation process. For 25 km altitude it takes about 1.5 days and decreases with altitude, so that at 16 km a time of several hours is needed. In comparison the time scale for adding an H_2SO_4 molecule to a CN is approximately 3 h (Arnold, 1982). Note these values are very variable depending on H_2SO_4 concentration and aerosol particle distribution.

We have to consider that some additional physical processes can strongly influence the aerosol formation mechanism for example: additional H_2SO_4 sources, such as hydroxyl radicals (OH) that seem to appear during GLE events (Krieger and Arnold, 1994); decrease of

stratospheric temperatures during SPE events (Kodama *et al.*, 1992); sudden stratospheric warming and subsequent sudden cooling that can increase level of H_2SO_4 supersaturation etc. According to the model suggested the observed $R(H)$ enhancement was caused by sulfate aerosol particles with radii $r \geq 0.69 \mu\text{m}$. It is known that liquid sulfate aerosols of such sizes can exist in the stratosphere under temperatures which are lower than type-1 PSC creation temperatures (Dye *et al.*, 1992; Larsen, 1994). On the other hand, one can not exclude the type-1 PSC creation because the temperatures are quite low enough for it to occur (see Fig. 4). In this case the sulfate aerosol particles may be condensation nuclei for type-1 PSCs (Tolbert, 1994). Some other physical mechanisms are also possible. For example, according to Tinsley *et al.* (1989) and Tinsley and Heelis (1993), ionization charges super-cooled droplets in the stratosphere causing droplet growth and nucleation process intensification.

Results of the work show that GLE events seem to be able to trigger aerosol creation at high latitudes. Some experimental facts confirm the probable connection between solar proton events and aerosol concentration increase (Hofman and Rosen, 1983; Pudovkin and Babushkina, 1992; Roldugin and Vashenyuk, 1994). In Shumilov *et al.* (1992; 1995) reported an ozone total content decrease during a GLE event and Tinsley *et al.* (1993) have demonstrated the connection between the vorticity atmospheric index (VAI) and galactic cosmic ray variations which also indirectly confirm the reality of such connections.

4 Conclusions

1. A considerable increase of aerosol concentration up to 50% at 20 km level after GLE event on 16 February 1984 was documented by lidar measurements at Verkhnetulomski observatory.
2. The profile of CN calculated using ion nucleation mechanism is presented. By modelling the growth of CN through coagulation process it is shown that theoretical and experimental aerosol altitude profiles coincide at some altitude ranges.
3. The results show the very important additional consequence that the transparency of the atmosphere can be greatly decreased after GLE events.

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